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## Corral Framework: Trustworthy and Fully Functional Data Intensive Parallel Astronomical Pipelines

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#### Abstract

Data processing pipelines represent an important slice of the astronomical sufficiency inbrary that include chains of processes that transform raw data into valuable information via data reduction and analysis. In this work we present Corral, a Python framework for astronomical pipeline generation. Corral features a Model-View-Control. design pattern on top of an SQL Relational Database capable of handling: custom data models; processing stages; and communication alerts, and also provides automatic quality and structural metrics based on unit testing. The Model-View-Controller processing concept separation between the user logic and the data models, delivering at the same time multi-processing and distributed computing capabilities. Corral represents an improvement over commonly found data processing pipelines in Astron and the same time provides a broad measure of quality over the created pipeline. Corral and working examples of pipelines that use it are available to the community at https. // github.com/toros-astro.

Keywords: Astroinformatics, Astronomical Pipeline, Softy, re and its engineering: Multiprocessing; Design Patterns

#### 1. Introduction

The development of modern ground-based and space-born telescopes, covering all observable window. in ne electromagnetic spectrum, and an ever increasing variab. it interest via time-domain astronomy have raised the processity for large databases of astronomical observations. The and unit of data to be processed has been steadily increasing, imposing higher demands over: quality; storage need: and analysis tools. This phenomenon is a manifestation of the dee transformation that Astronomy is going through, alchie view the development of new technologies in the Big Data core. In this context, new automatic data analysis technique charge data the preferred solution to the so-called "dat tsunan." (Cavuoti, 2013).

The development of an intermatio processing pipeline is a natural consequence of science projects involving the acquisition of data and its post rior and vsis. Some examples of these data intensive projects in hide the Dark Energy Survey Data Management Syster (Mohr et al., 2008), designed to exploit a camera with 74 CCD, at the Blanco telescope to study the nature of cosmic acceleration; the Infrared Processing and Analysis Center (Masci et al., 2016), a near real-time transient-source

discovery engine for the intermediate Palomar Transient Factory (iPTF Kulkarni, 2013); and the Pan-STARRS PS1 Image Processing Pipeline (Magnier et al., 2006), performing the image processing and analysis for the Pan-STARRS PS1 prototype telescope data and making the results available to other systems within Pan-STARRS and Vista survey pipeline that includes VIRCAM, a 16 CCD nearIR camera for the VISTA Data flow system Emerson et al. (2004). In fact, the implementation of pipelines in Astronomy is a common task to the construction of surveys (e.g. Marx and Reyes, 2015; Hughes et al., 2016; Hadjiyska et al., 2013), and it is even used to operate telescopes remotely, as described in Kubánek et al. (2010). Standard tools for pipeline generation have already been developed and can be found in the literature. Some examples are Luigi<sup>1</sup>, which implements a method for the creation of distributive pipelines; OPUS (Rose et al., 1995), conceived by the Space Telescope Science Institute; and more recently Kira (Zhang et al., 2016), a distributed tool focused on astronomical image analysis. In the experimental sciences, collecting, pre-processing and storing data are common recurring patterns regardless of the science field or the nature of the experiment. This means that pipelines are in some sense re-written repeatedly. A more efficient approach would exploit existing resources to build new tools and perform new tasks, taking advantage of established procedures

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<sup>&</sup>lt;sup>1</sup>Luigi: https://luigi.readthedocs.io/

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